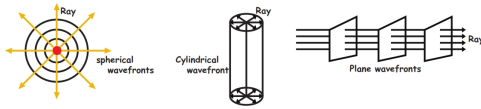


# WAVE OPTICS

## Wave Front

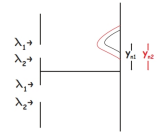


Point light source → spherical wavefront  
 Linear light Source → cylindrical wavefront  
 Source at infinity → Plane wave front

## Huygen's principle

- Each point on a wavefront acts as a fresh source of new disturbance, called secondary waves or wavelets. The secondary wavelets spread out in all directions with the speed of light in the given medium.
- A common envelope or common tangent to these secondary wavelets at any later time gives secondary wavefront at that time

## Overlapping



Let  $n_1 \lambda_1$  max of  $\lambda_1$  wavelength overlaps with  $n_2 \lambda_2$  max of  $\lambda_2$  wavelength

$$y_{n1} = y_{n2}$$

$$\frac{n_1 D \lambda_1}{d} = \frac{n_2 D \lambda_2}{d}$$

$$n_1 \lambda_1 = n_2 \lambda_2$$

- As we move further away, then overlapping of colours increases if white light is used
- At larger distance, all colours again overlap to give white light pattern

	Incident wavefront	Reflected wavefront
<b>Concave Mirror</b> 	Plane wavefront	Spherical converging wavefront
<b>Convex Mirror</b> 	Plane wavefront	Spherical diverging wavefront
	Incident wavefront	Refracted wavefront
<b>Convex Lens</b> 	Plane wavefront	Spherical converging wavefront
<b>Concave Lens</b> 	Plane wavefront	Spherical diverging wavefront

## Phase Difference & Path Difference

$$\Phi = \frac{2\pi}{\lambda} \Delta x$$

## Phase Difference & Time Difference

$$\Phi = \frac{2\pi}{T} \Delta t$$

## Resultant Amplitude

$$y_1 = A_1 \sin \omega t \text{ and}$$

$$y_2 = A_2 \sin (\omega t + \Phi)$$

$$\text{Resultant } A = \sqrt{A_1^2 + A_2^2 + 2A_1 A_2 \cos \Phi}$$

- $\cos \Phi = 1 \Rightarrow A = A_{\max} = \sqrt{(A_1 + A_2)^2} = A_1 + A_2$
- $\cos \Phi = -1 \Rightarrow A = A_{\min} = \sqrt{(A_1 - A_2)^2} = A_1 - A_2$

$$\frac{A_{\max}}{A_{\min}} = \frac{A_1 + A_2}{A_1 - A_2}$$

- Intensity  $\propto$  (amplitude)<sup>2</sup>

## Resultant Intensity

$$\text{We have, } I = I_1 + I_2 + 2\sqrt{I_1 I_2} \cos \Phi$$

$$\cdot \cos \Phi = 1 \Rightarrow I = I_{\max}$$

$$I_{\max} = I_1 + I_2 + 2\sqrt{I_1 I_2}$$

$$I_{\max} = (\sqrt{I_1} + \sqrt{I_2})^2$$

$$\cdot \cos \Phi = -1 \Rightarrow I = I_{\min}$$

$$I_{\min} = (\sqrt{I_1} - \sqrt{I_2})^2$$

$$I_{\max} = \frac{(\sqrt{I_1} + \sqrt{I_2})^2}{(\sqrt{I_1} - \sqrt{I_2})^2}$$

$$I_{\max} \propto A_{\max}^2 \text{ \& } I_{\min} \propto A_{\min}^2$$

$$I_{\max} = \frac{A_{\max}^2}{A_{\min}^2} = \frac{(A_1 + A_2)^2}{(A_1 - A_2)^2}$$

$$I_{\min} = \frac{A_{\min}^2}{A_{\max}^2} = \frac{(A_1 - A_2)^2}{(A_1 + A_2)^2}$$

$$\text{If } I_1 = I_2 = I_0$$

$$\Rightarrow I = 4I_0 \cos^2 \frac{\Phi}{2}$$

## Young's Double-slit experiment (YDSE)

$$\text{Path difference } \Delta X = \frac{y_n d}{D}$$

$$\text{In general } \Delta X = \frac{y_n d}{D}$$

Distance of Minima and Maxima from Central maximum

Maxima

Minima

$$y_n = \frac{n D \lambda}{d} \quad n = 0, 1, 2, \dots \quad y_n = (2n-1) \frac{D \lambda}{2d} \quad n = 1, 2, \dots$$

## YDSE in Liquid

When YDSE setup is immersed in a liquid, there is change in wavelength

$$n = \frac{c}{v} = \frac{v \lambda}{v \lambda'} = \frac{\lambda}{\lambda'} \quad n \rightarrow \text{refractive index}$$

$$\lambda_{\text{medium}} = \lambda' = \frac{\lambda}{n} \quad \text{or } \lambda' = \frac{\lambda}{\mu} \quad \mu \rightarrow \text{refractive index}$$

$$\text{In air } y_n = n \frac{D \lambda}{d}$$

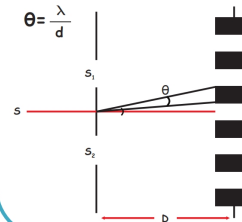
$$\text{In medium } y'_n = \frac{n D \lambda'}{d} = \frac{n D \lambda}{d \mu}$$

$$\text{Fringe width in air } \beta = \frac{D \lambda}{d}$$

$$\text{In medium } \beta' = \frac{D \lambda'}{\mu d} \Rightarrow \beta_{\text{med}} = \frac{\beta_{\text{air}}}{\mu}$$

$$\Rightarrow \beta_{\text{med}} < \beta_{\text{air}}$$

## Angular fringe width



## Constructive interference

Phase difference at the point of observation  
 $\Phi = 0^\circ$  or  $2n\pi$ ,  $n = 0, 1, 2, \dots$

Also,  $\Delta x = n \lambda$ ,  $n = 0, 1, 2, \dots$

Resultant intensity at the point of observation is maximum

$$I = I_{\max} = (\sqrt{I_1} + \sqrt{I_2})^2$$

## Destructive interference

$\Phi = 180^\circ$   
 or  $\Phi = (2n-1)\pi$ ;  $n = 1, 2, \dots$

Also,  $\Delta x = (2n-1) \frac{\lambda}{2}$ ,  $n = 1, 2, 3, \dots$

Resultant intensity at the point of observation will be minimum

$$I = I_{\min} = (\sqrt{I_1} - \sqrt{I_2})^2$$

## Intensity at any point on screen

For all maxima  $I = 4I_0$  (If  $I_1 = I_2 = I_0$ )

For all minima,  $I = 0$

Note:

$$\text{Fringe visibility } V = \frac{I_{\max} - I_{\min}}{I_{\max} + I_{\min}}$$

## Introduction Of Thin Transparent Sheet in YDSE

Optical path length and geometrical path length

$$\text{Refractive index } \mu = \frac{c}{v}$$

$$\mu = \frac{v \lambda}{v \lambda_m} \Rightarrow \lambda_m = \frac{\lambda}{\mu}$$

Time taken by light to travel x length in medium,

$$t = \frac{x}{c/\mu} = \frac{\mu x}{c}$$

Distance travelled by light in vacuum in same time = optical path length

$$\text{Optical Path Length (OPL)} = \text{velocity} \times \text{time} = c \times \frac{\mu x}{c} = \mu x$$

If Geometrical Path Length (GPL) = x, then OPL =  $\mu x$ , where  $\mu$  is the refractive index of the medium

## Fringe Width or Band width ( $\beta$ )

$$\beta_{\text{dark}} = \frac{D \lambda}{d}$$

$$\beta_{\text{bright}} = \frac{D \lambda}{d}$$

$$\text{For interference pattern } \beta_{\text{dark}} = \beta_{\text{bright}} = \frac{D \lambda}{d}$$

### Introduction of thin transparent sheet

Path difference  $\Delta x = s_2P - s_1P$

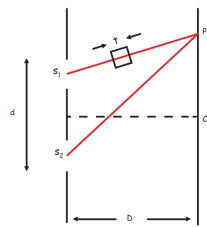
Additional path difference =  $(\mu - 1)t$

Geometrical path difference before inserting sheet,  $\Delta x = \frac{y d}{D}$

$$y = \frac{D}{d} \Delta x$$

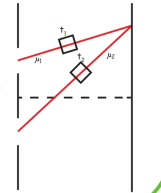
After introducing sheet,  $y' = \frac{D}{d} [\Delta x + (\mu - 1)t]$

Shift  $S = y' - y = \frac{D}{d} (\mu - 1)t$



If two plates are introduced,

Shift  $S = [(\mu_1 - 1)t_1 - (\mu_2 - 1)t_2] \frac{D}{d}$



**Interference of reflected light :**

For normal incidence  $r = 0$  so,  $2\mu t = (2n - 1) \frac{\lambda}{2}$

**Interference of refracted light :**

For normal incidence  $2\mu t = n\lambda$

### Single Slit Diffraction

Path difference =  $\Delta x = d \sin \theta$

• Formation of first secondary minima

Path difference =  $\frac{\lambda}{2}$

• Formation of 2nd secondary minima

Path difference =  $2\lambda$

• Formation of n<sup>th</sup> secondary minima

$d \sin \theta_n = n\lambda$

$n = 1, 2, 3, \dots$

### First secondary maxima

But the intensity of 1<sup>st</sup> secondary maxima is lower than central maximum

N<sup>th</sup> secondary Maxima

$x = (2n + 1) \frac{\lambda}{2}$

$n = 1, 2, 3, \dots$

$d \sin \theta_n = (2n + 1) \frac{\lambda}{2}$

Ratio of intensities of central maxima and secondary maxima

$1 : \frac{1}{2} : \frac{1}{61} : \frac{1}{121} : \dots$

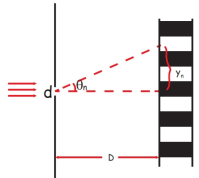
### Distance of N<sup>th</sup> secondary maxima from CM

$\theta_n = (2n + 1) \frac{\lambda}{2d}$

$n = 1, 2, 3, \dots$

$y = (2n + 1) \frac{D\lambda}{d}$

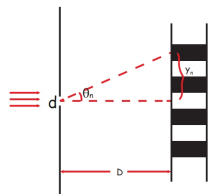
$n = 1, 2, 3, \dots$



### Distance of N<sup>th</sup> secondary minima from CM

$y = \frac{nD\lambda}{d}$

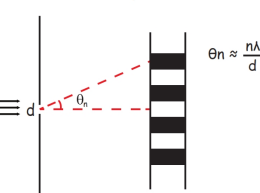
$n = 1, 2, 3, \dots$



# WAVE OPTICS

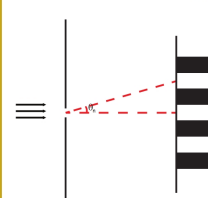
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### Angular position of N<sup>th</sup> secondary minima

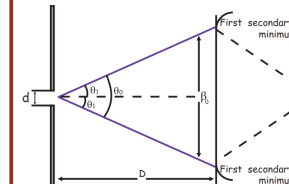


### Angular position of N<sup>th</sup> secondary Maxima

$\theta_n \approx (2n + 1) \frac{\lambda}{2d}$



### Angular width and linear width of central maximum



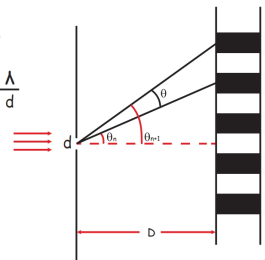
$\beta_0 = \theta_0 D = \frac{2D\lambda}{d}$

### Angular width and linear width of secondary maxima

Angular position of n<sup>th</sup> minimum,  $\theta_n = n \frac{\lambda}{d}$

Angular position of n+1<sup>th</sup> minimum,  $\theta_{n+1} = (n+1) \frac{\lambda}{d}$

Linear width,  $\beta = \frac{D\lambda}{d}$



### Angular width and Linear width of secondary minima

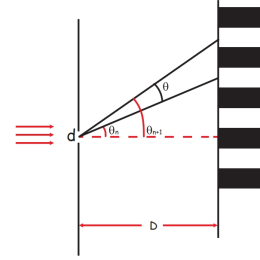
Angular position of n<sup>th</sup> maximum  $\theta_n = (2n + 1) \frac{\lambda}{2d}$

Angular position of (n+1)<sup>th</sup> maximum  $\theta_{n+1} = (2(n+1) + 1) \frac{\lambda}{2d}$

$\theta = \frac{\lambda}{d}$

Linear width  $\beta = \theta D$

$\beta = \frac{D\lambda}{d}$



### Validity of Ray Optics: Fresnel's Distance

$Z_F = \frac{d^2}{\lambda}$

### Resolving Power (R.P) of a microscope

$\frac{1}{d} = \frac{2n \sin \theta}{\lambda}$

### Resolving Power

$R.P = \frac{1}{\text{limit of resolution}}$

### Resolving power of a telescope

$R.P = \frac{1}{d\theta} = \frac{D}{1.22 \lambda}$

### Law of Malus

$I \propto \cos^2 \theta$

$I = I_0 \cos^2 \theta$

When  $\theta = 0^\circ$  or  $180^\circ$ ,  $\cos \theta = \pm 1 \Rightarrow I = I_0$

When  $\theta = 90^\circ$ ,  $\cos \theta = 0 \Rightarrow I = 0$

### Polarisation by Reflection

Brewster found that at the polarising angle, the reflected and transmitted rays are perpendicular to each other

$n = \tan i_p$

This is Brewster's Law

